# OPERATOR FUNCTION BY CHEBYSHEV FIT 

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#### Abstract

We explain the concept of orthogonal polynomials and specialize to Chebyshev polynomials. We then explain how to use this expansion for computing an operator function on a Hilbert-space function. As applications we discuss how to efficiently computer the free energy and the entropy of a particle in temperature $T$.


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## 1. Introduction

In many cases in Quantum Mechanics we need to apply a function $f(\hat{H})$ of an operator $\hat{H}$ to a wavefunction $\psi_{0}$. For example, when solving the time-dependent Schrodinger, starting from a given initial state $\psi(0)=\psi_{0}$ at time $t=0$ :

$$
\begin{equation*}
i \hbar \dot{\psi}(t)=\hat{H} \psi(t) \tag{1.1}
\end{equation*}
$$

where $\hbar$ is Planck's constant. The solution at time $t$ is:

$$
\begin{equation*}
\psi(t)=e^{-i \hat{H} t / \hbar} \psi_{0} \tag{1.2}
\end{equation*}
$$

Here the operator function is based on the imaginary exponential function $f_{t}(x)=$ $e^{-i x t / \hbar}$ and has time as a parameter $t$.

Another example is the partition function which is defined as a trace of an operator function:

$$
\begin{equation*}
Z=\operatorname{tr}\left[e^{-\beta \hat{H}}\right] \tag{1.3}
\end{equation*}
$$

where now the operator function is the real exponential function $f_{\beta}(x)=e^{-\beta x}$ and has the inverse temperature as a parameter $\beta$.

For non-interacting fermions the number of fermions as a function of temperature and chemical potential $\mu$ :

$$
\begin{equation*}
N(\beta, \mu)=\operatorname{tr}\left[\frac{2}{1+e^{\beta(\hat{H}-\mu)}}\right] \tag{1.4}
\end{equation*}
$$

Here we will describe how such an operator function can be computed using orthogonal polynomials.

## 2. Orthogonal Polynomials

Definition 1. A set of polynomials $p_{n}(x) n=0,1,2, \ldots$ is called an orthogonal set if $p_{n}(x)$ is of degree $n$ and if

$$
\begin{equation*}
\int_{a}^{b} p_{n}(x) p_{m}(x) w(x) d x=c_{n} \delta_{n m} \tag{2.1}
\end{equation*}
$$

We denote the coefficient of $x^{n}$ (the leading coefficient) in $p_{n}(x)$ by $A_{n} . w(x)>0$ is the weight function so clearly the norm $c_{n}$ is positive.

Example 2. Examples of orthogonal polynomials are the Legendre polynomials, orthogonal in the interval $[-1,1]$. These are obtained by a Gram Schmidt orthogonalization of the set $\left\{1, x, x^{2}, \ldots\right\}$. Remember that the Gram Schmidt procedure for orthogonalization of a set of vectors $\alpha_{n}$ is the set $\beta_{0}=\alpha_{0}$ and

$$
\begin{equation*}
\beta_{n}=\alpha_{n}-\sum_{k=0}^{n-1} \frac{\beta_{k}\left\langle\beta_{k} \mid \alpha_{n}\right\rangle}{\left\langle\beta_{k} \mid \beta_{k}\right\rangle} . \tag{2.2}
\end{equation*}
$$

In our case, take $p_{0}(x)=1$. Then, since $\left\langle p_{0} \mid x\right\rangle=0$

$$
\begin{equation*}
p_{1}(x)=x . \tag{2.3}
\end{equation*}
$$

Next,

$$
\begin{align*}
p_{2}(x) & =x^{2}-p_{1}(x) \frac{\left\langle p_{1} \mid x^{2}\right\rangle}{\left\langle p_{1} \mid p_{1}\right\rangle}-p_{0}(x) \frac{\left\langle p_{0} \mid x^{2}\right\rangle}{\left\langle p_{0} \mid p_{0}\right\rangle}  \tag{2.4}\\
& =x^{2}-\frac{2}{3} \tag{2.5}
\end{align*}
$$

and

$$
\begin{equation*}
p_{3}(x)=x^{3}-p_{1}(x) \frac{\left\langle p_{1} \mid x^{3}\right\rangle}{\left\langle p_{1} \mid p_{1}\right\rangle}=x^{3}-\frac{3}{5} x \tag{2.6}
\end{equation*}
$$

Where we used: $\int_{-1}^{1} x^{2 n} d x=\frac{2}{2 n+1}$ and $\int_{-1}^{1} x^{2 k} d x=0$. Continuing this way we get $p_{4}(x)=x^{4}-\frac{6}{7} x^{2}+\frac{3}{35}$ etc.
Lemma 3. The polynomial $x^{n}$ is a linear combination of the orthogonal set of polynomials $p_{m}(x)$ where $m \leq n$.
Proof. Using induction. For $n=0$ it is trivial. Now suppose it is true for $n>0$ then for $n+1$ we have that $x^{n+1}-\frac{1}{A_{n+1}} p_{n+1}(x)$ is a polynomial of degree $n$ and thus equal to $\sum_{m=0}^{n} b_{m} p_{m}(x)$. Thus, $x^{n+1}=\frac{1}{A_{n+1}} p_{n+1}(x)-\sum_{m=0}^{n} b_{m} p_{m}(x)$ i.e. it is a linear combination of $p_{m}(x)$ 's.
Corollary 4. Within a set of orthogonal polynomials, $p_{n}$ is orthogonal to all mdegree polynomials $q_{m}(x)$ where $m<n$.

Proof. For $m<n p_{m}(x)$ is a linear combination of $p_{k}(x) k \leq m$, all of which are orthogonal to $p_{n}(x)$ by definition.
Corollary 5. Any m-degree polynomial $q(x)$ can be written as $q(x)=\sum_{n=0}^{m} b_{n} p_{n}(x)$ where $b_{n}=c_{n}^{-1} \int q(x) p_{n}(x) w(x) d x$.
Definition 6. A root of a polynomial $p(x)$ is a point of sign change. Specifically, a point $t$ for which $p(x) p(y)<1$ for all $x<t<y$ in an infinitesimal interval around $t$. Naturally $p(t)=0$.

Theorem 7. All $n$ roots of $p_{n}(x)$ are real, distinct and appear inside the interval $[a, b]$.
Proof. Denote by $N<=n$ the number of sign changes of $p_{n}(x)$ in $[a, b]$. If $N=0$ then define $q_{0}(x)=1$ and if $N>0$ define polynomial $q_{N}(x)=\Pi_{i=1}^{N}\left(x-t_{i}\right)$, where $t_{i}$ are the $N$ roots of $p_{n}(x)$ in the interval. $p_{n}(x) q_{N}(x)$ does not change sign even once in the interval and therefore $\int_{a}^{b} p_{n}(x) q_{N}(x) w(x) d x \neq 0$ i.e. $p_{n}$ and $q_{N}$ are non-orthogonal. This can only be true if $N=n$ since $p_{n}$ is orthogonal to all polynomials of degree $N<n$. Hence $p_{n}(x)$ changes sign $n$ times in the interval, which means that all its roots are distinct and real.

## 3. Quadrature as approximations to integrals

Definition 8. A rule $\sum_{n=1}^{N} a_{n} f\left(x_{n}\right)$ is called a quadrature formula for $f(x)$ with $N$ nodes $x_{n}$ or a $N$-quadrature; $a_{n}$ are called the weights. We use quadratures to approximate integrals $\int_{a}^{b} f(x) w(x) d x$. A $N$-quadrature is called "k-exact" if it is exact for all $f(x)=q_{m}(x)$ polynomials of degrees $m \leq k$.

Theorem 9. A $N$-quadrature cannot be $2 N$-exact.
Proof. By counter example: $f(x)=\Pi_{n=1}^{N}\left(x-x_{n}\right)^{2}$ is a non-negative $2 N$-degree polynomial. Its integral is positive while the quadrature formula yields zero.
Definition 10. A $N$-quadrature which is $(2 N-1)$-exact is called a Gaussian N -quadrature.

Theorem 11. A $N$-quadrature which is $(N-1)$-exact is also $(2 N-1)$-exact if and only if the nodes $x_{n}$ are the roots of $p_{N}(x)$.
Proof. Follow these:
(1) If the nodes $x_{n}$ are the roots of $p_{N}(x)$ then the quadrature is $(2 N-1)$ exact: Any polynomial $s_{k}(x)$ of degree $k<2 N$ can be written as $s_{k}(x)=$ $q(x) p_{N}(x)+r(x)$ where both $q(x)$ and $r(x)$ are polynomials of degree less than $N$. Now $\int s_{k}(x) w(x) d x=\int r(x) w(x) d x$ since $p_{N}(x)$ is orthogonal to $q(x)$. Finally, since $q\left(x_{n}\right) p_{N}\left(x_{n}\right)=0$ and the quadrature is $(N-1)$ exact we have: $\int s_{k}(x) w(x) d x=\sum_{n=1}^{N} a_{n} r\left(x_{n}\right)=\sum_{n=1}^{N} a_{n} s_{k}\left(x_{n}\right)$, hence it is also $2 N-1$ exact.
(2) In a $N$-quadrature which is $(2 N-1)$-exact the nodes $x_{n}$ must be the roots of $p_{N}(x)$ : Let $q_{N}(x)=\Pi_{n=1}^{N}\left(x-x_{n}\right)$ then for $k<N: \int q_{N}(x) p_{k}(x) w(x) d x=$ $\sum_{n=1}^{N} a_{n} q_{N}\left(x_{n}\right) p_{k}\left(x_{n}\right)=0$ since then the quadrature is exact. Thus $q_{N}(x)$ is a $N$ degree polynomial which is orthogonal to all $p_{k}(x)$ i.e. it must be a constant times $p_{N}(x)$.

Corollary 12. The weights are obtained from solving the $N$ linear equations with $N$ unknowns

$$
\begin{equation*}
\sum_{n=1}^{N} a_{n} p_{k}\left(x_{n}\right)=\frac{c_{0}}{p_{0}(0)} \delta_{k 0}, k=0, \ldots, K \tag{3.1}
\end{equation*}
$$

Theorem 13. For a Gaussian quadrature $\sum_{n=1}^{N} a_{n} p_{k}\left(x_{n}\right) p_{j}\left(x_{n}\right)=c_{k} \delta_{k j}, k, j=$ $0, \ldots, K$

## 4. Polynomial fit approximation to functions

Given a function $f(x)$, an interval $[a, b]$ and a weight $w(x)$ and a set of polynomials of degree $k=0,1, \ldots p_{k}(x)$ we ask what is the best fit we can obtain for $f(x)$

$$
\begin{equation*}
f(x) \approx \sum_{n=0}^{N-1} \alpha_{n} p_{n}(x) \tag{4.1}
\end{equation*}
$$

we can obtain for $f(x)$ using these polynomials, so that the difference is minimal in the sense

$$
\begin{equation*}
J=\int_{a}^{b}\left(f(x)-\sum_{n=0}^{N-1} \alpha_{n} p_{n}(x)\right)^{2} w(x) d x \tag{4.2}
\end{equation*}
$$

We can solve this by taking derivatives with respect to $\alpha_{k}$ :

$$
\begin{equation*}
0=\frac{\partial J}{\partial \alpha_{k}}=-2 \int_{a}^{b}\left(f(x)-\sum_{n=0}^{N-1} \alpha_{n} p_{n}(x)\right) p_{k}(x) w(x) d x \tag{4.3}
\end{equation*}
$$

which can be rearranged to read:

$$
\begin{equation*}
\sum_{n=0}^{N-1} \alpha_{n} \int_{a}^{b} p_{n}(x) p_{k}(x) w(x) d x=\int_{a}^{b} f(x) p_{k}(x) w(x) d x \tag{4.4}
\end{equation*}
$$

we see that, if we choose the polynomials $p_{n}(x)$ as the orthogonal set with respect to $w(x)$ we obtain a simple expression :

$$
\begin{equation*}
\alpha_{k}=\frac{1}{c_{k}} \int_{a}^{b} f(x) p_{k}(x) w(x) d x \tag{4.5}
\end{equation*}
$$

Hence we have a systematic way of generating optimal approximations to functions. If we know how the perform the integrals of Eq. 4.5 we can use the obtained $\alpha_{k}$ for approximating the function of an operator having its eigenvalues within the interval $[a, b]$.

Now, if we cannot use exact integration, we can use Gauss integration:

$$
\begin{equation*}
\alpha_{k}=\frac{1}{c_{k}} \sum_{n=1}^{N^{\prime}} f\left(t_{n}\right) p_{k}\left(t_{n}\right) w_{n}, \quad k=0, \ldots, N-1 \tag{4.6}
\end{equation*}
$$

One can choose $N^{\prime}=N$ or even $N^{\prime}=2 N$ for higher precision.

## 5. Chebyshev Polynomials

The Chebyshev polynomials $T_{n}(x)$ are a family of polynomials, defined through trigonometric functions. For $|x| \leq 1$ we set $x=\cos \theta$ and define

$$
\begin{equation*}
T_{n}(x)=T_{n}(\cos \theta)=\cos n \theta \tag{5.1}
\end{equation*}
$$

Is this really a polynomial?? First we note:

$$
\begin{aligned}
& T_{0}(x)=1 \\
& T_{1}(x)=x
\end{aligned}
$$

Now, using the trigonometric relation:

$$
\cos (n+1) \theta=2 \cos \theta \cos n \theta-\cos (n-1) \theta
$$

It is straightforward to unveil the polynomial nature: $T_{n+1}(x)$ can be defined in terms of lower degree T's :

$$
\begin{equation*}
T_{n+1}(x)=2 x T_{n}(x)-T_{n-1}(x) \tag{5.2}
\end{equation*}
$$

Thus the Chebyshev polynomials can be generated, for example

$$
\begin{aligned}
& T_{2}(x)=2 x^{2}-1 \\
& T_{3}(x)=4 x^{3}-3 x \\
& T_{4}(x)=8 x^{4}-8 x^{2}+1
\end{aligned}
$$

The orthogonality of the Chebyshev polynomials over the interval $[-1,1]$ derives from:

$$
\begin{equation*}
\frac{2}{\pi} \int_{0}^{\pi} \cos (n \theta) \cos (m \theta) d \theta=\left(\delta_{n 0}+1\right) \delta_{n m} \tag{5.3}
\end{equation*}
$$

and by making the change of variable $x=\cos \theta$, we obtain:

$$
\begin{equation*}
\frac{2}{\pi} \int_{-1}^{1} \frac{T_{n}(x) T_{m}(x)}{\sqrt{1-x^{2}}} d x=\delta_{n m}\left(\delta_{n 0}+1\right) \tag{5.4}
\end{equation*}
$$

Hence, the weight $w(x)=\frac{2}{\pi} \frac{1}{\sqrt{1-x^{2}}}$ emerges.
The roots of the Chebyshev polynomial $T_{N}(x)$ are derived from the zeros of the cosine $\cos N \theta=0$ which are

$$
\begin{equation*}
\theta_{n}^{(N)}=\frac{\pi}{N}\left(n-\frac{1}{2}\right) \tag{5.5}
\end{equation*}
$$

where $n$ is any integer. Hence, the Gaussian-Chebyshev sampling points, based on the $T_{N}(x)$ the are

$$
\begin{equation*}
x_{n}^{(N)}=\cos \theta_{n}^{(N)}, \mathrm{n}=1, \ldots N \tag{5.6}
\end{equation*}
$$

The $N$ quadrature weights weights $a_{n}$ are determined from Eq. 3.1 one gets: $\sum_{n=1}^{N} a_{n} T_{k}\left(x_{n}^{(N)}\right)=2 \delta_{k 0}$ or

$$
\begin{equation*}
\sum_{n=1}^{N} a_{n}^{(N)} \cos \left(k \theta_{n}^{(N)}\right)=2 \delta_{k 0}, k=0, \ldots, N-1 \tag{5.7}
\end{equation*}
$$

Due to the fact that This is a set of $N$ equations and $N$ unknowns. We note that for $k \neq 0$

$$
\begin{align*}
\sum_{n=1}^{N} \cos \left(k \theta_{n}^{(N)}\right) & =R e \sum_{n=0}^{N-1} e^{i k \frac{1}{N}\left(\frac{\pi}{2}+n \pi\right)}  \tag{5.8}\\
& =R e\left[e^{\left.i k \frac{\pi}{2 N} \frac{(-)^{k}-1}{2 i \sin \frac{k \pi}{2 N}}\right]}\right.  \tag{5.9}\\
& =0 \tag{5.10}
\end{align*}
$$

this gives:

$$
\begin{equation*}
a_{n}^{(N)}=\frac{2}{N} \tag{5.11}
\end{equation*}
$$

The quadrature is then :

$$
\begin{equation*}
\frac{2}{\pi} \int_{-1}^{1} \frac{f(x)}{\sqrt{1-x^{2}}} d x \approx \frac{2}{N} \sum_{n=1}^{N} f\left(\cos \theta_{n}^{(N)}\right) \tag{5.12}
\end{equation*}
$$

## 6. Chebyshev approximation

Further reading: [1]
Suppose we have a set of functions $f(x ; p)$ where $x$ is a variable in the interval $[-1,1]$ and $p$ is some set of parameters. We can use Chebyshev polynomials and write,

$$
\begin{equation*}
f(x ; p) \approx \sum_{k=0}^{N-2} b_{k}(p) T_{k}(x) \tag{6.1}
\end{equation*}
$$

where due to Eq. 5.4:

$$
\begin{equation*}
b_{k}(p)=\frac{1}{1+\delta_{k 0}} \frac{2}{\pi} \int_{-1}^{1} \frac{T_{k}(x) f(x ; p)}{\sqrt{1-x^{2}}} d x \tag{6.2}
\end{equation*}
$$

Then, using the Gaussian quadrature of with $N$ terms in Eq. (5.12) we estimate $b_{k}, k=0, \ldots, N-2$ as:

$$
\begin{equation*}
b_{k}(p)=\frac{1}{1+\delta_{n 0}} \frac{2}{N} \sum_{n=1}^{N} \cos \left(k \theta_{n}^{(N)}\right) f\left(\cos \theta_{n}^{(N)} ; p\right) \tag{6.3}
\end{equation*}
$$

Note: The constant $b_{k}(p)$ can be calculated efficiently using a FFT. Denoting $f_{n}(p)=f\left(\cos \theta_{n}^{(N)} ; p\right)$, for real functions $f(x)$ We write Eq. (6.3) as:

$$
\begin{equation*}
b_{k}(p)=\frac{1}{1+\delta_{n 0}} \frac{2}{N} R e\left\{e^{i \frac{\pi}{2 N} k} \sum_{n=0}^{2 N-1} e^{i \frac{2 \pi}{2 N} n k} \phi_{n}(p)\right\}, k=0, \ldots, N-2 \tag{6.4}
\end{equation*}
$$

Where the sum can be done using a FFT, by setting

$$
\phi_{n}=\left\{\begin{array}{cc}
f_{n+1} & 0 \leq n<N  \tag{6.5}\\
0 & \\
\hline \leq n<2 N
\end{array}\right.
$$

so the sum in the last term is a FFT of $f_{n}(p)$ (for each value of parameters $p$ an FFT needs to be performed) and all the $b$ 's for this set of parameters are obtained.

If $f(x)=f^{r}(x)+i f^{i}(x)$, is complex then you don't have to separate to real and imaginary parts. Just do the following FFT (note the :

$$
\begin{equation*}
b_{k}(p)=\frac{1}{1+\delta_{k 0}} \frac{1}{N}\left\{e^{i \frac{\pi}{2 N} k} \sum_{n=0}^{2 N-1} e^{i \frac{2 \pi}{2 N} n k} f_{n+1}(p)\right\}, k=0, \ldots, N-1 . \tag{6.6}
\end{equation*}
$$

This works well for both the real and the complex cases and for real functions gives exactly the same coefficients as Eq. 6.4.

## 7. CHEBYSHEV APPROXIMATION FOR A OPERATOR FUNCTION

Based on ref.[2].
We now show how to apply $F_{p}(\hat{H})$ on a given function $\psi$. Given the operator $\hat{H}$ we first need to estimate the interval $\left[E_{\min }, E_{\max }\right]$ which contains all of $\hat{H}$ 's eigenvalues. Then we define a function: $b-a=E_{\text {min }}, b+a=E_{\text {max }}$

$$
\begin{equation*}
f_{p}(x)=F_{p}(x \Delta E+\bar{E}) \tag{7.1}
\end{equation*}
$$

where:

$$
\begin{equation*}
\bar{E}=\frac{E_{\max }+E_{\min }}{2}, \Delta E=\frac{E_{\max }-E_{\min }}{2} \tag{7.2}
\end{equation*}
$$

Now find the Chebyshev approximation for $f(x)$ :

$$
\begin{equation*}
f_{p}(x)=\sum_{k=0}^{K} b_{k}(p) T_{k}(x) \tag{7.3}
\end{equation*}
$$

Then, here's how to apply $F_{p}(\hat{H})$ on any given $\psi$ :

$$
\begin{equation*}
F_{p}(\hat{H}) \psi=\sum_{k=0}^{K} b_{k}(p) \psi_{k} \tag{7.4}
\end{equation*}
$$

where:

$$
\begin{equation*}
\psi_{0}=\psi, \quad \psi_{1}=\hat{H}_{N} \psi_{0}, \quad \psi_{k+1}=2 \hat{H}_{N} \psi_{k}+\psi_{k-1} \tag{7.5}
\end{equation*}
$$

and

$$
\begin{equation*}
\hat{H}_{N}=\frac{\hat{H}-\bar{E}}{\Delta E} \tag{7.6}
\end{equation*}
$$

Note that the same set of $\psi$ 's can be used for many parameter s $p$, and that you only need to remember 4 functions.

## 8. Stochastic wave functions and traces using Chebyshev moments

The methods here are based on refs. [3, 4, 5][4][5].
If we want to calculate the trace of an operator function, (see for examples Eq. (1.3) or Eq. (1.4)):

$$
\begin{equation*}
A_{p}=\operatorname{tr}\left[F_{p}(\hat{H})\right] \tag{8.1}
\end{equation*}
$$

We first employ the stochastic trace formula, where we take a random wave function at grid-point $\mathbf{r}_{g}$,

$$
\begin{equation*}
\psi\left(\mathbf{r}_{g}\right)=\frac{e^{i \theta_{g}}}{\sqrt{\Delta V}} \tag{8.2}
\end{equation*}
$$

where $\Delta V$ is the grid spacing/area/volume. It can be shown that the average of the projection operator of such wave functions give the unit operator:

$$
\begin{equation*}
\langle\mid \psi\rangle\langle\psi \mid\rangle=\mathbf{I} \tag{8.3}
\end{equation*}
$$

where the angular brackets denote averaging with respect to $\theta$. Combining this with Eq. (8.1) we obtain:

$$
\begin{equation*}
\left.A_{p}=\operatorname{tr}\left[\langle\mid \psi\rangle\langle\psi \mid\rangle F_{p}(\hat{H})\right]=\left\langle\langle\psi| F_{p}(\hat{H}) \mid \psi\right\rangle\right\rangle \tag{8.4}
\end{equation*}
$$

Using the Chebyshev approximation of the function in Eq. (7.4) we find:

$$
\begin{equation*}
A_{p}=\sum_{k=0}^{K-1} b_{k}(p) M_{k} \tag{8.5}
\end{equation*}
$$

where:

$$
\begin{equation*}
M_{k}=\operatorname{tr}\left[T_{k}\left(\hat{H}_{N}\right)\right]=\left\langle\left\langle\psi \mid \psi_{k}\right\rangle\right\rangle \tag{8.6}
\end{equation*}
$$

is the $k$ Chebyshev moment. Note that the main numeric effort goes to calculation of the $K$ moments, which are just numbers. One can save half of the Hamiltonian applications by noticing that from the trigonometric relations $\cos (n+m) \theta+$ $\cos (n-m) \theta=2 \cos n \theta \cos m \theta$, we have, for $-1 \leq x \leq 1$ :

$$
\begin{equation*}
T_{n+m}(x)+T_{|n-m|}(x)=2 T_{n}(x) T_{m}(x) . \tag{8.7}
\end{equation*}
$$

Using this relation with $n=m$ (we omit $(x)$ ):

$$
\begin{equation*}
T_{2 n}=2 T_{n}^{2}-1 \tag{8.8}
\end{equation*}
$$

and with $n=m+1$ we find:

$$
\begin{equation*}
T_{2 n+1}=2 T_{n} T_{n+1}-T_{1} \tag{8.9}
\end{equation*}
$$

Hence, applying this for the operator $\hat{H}_{N}$ instead of $x$, for even $k$, i.e. $k=2 n$, we find:

$$
\begin{equation*}
M_{2 n}=2\left\langle\left\langle\psi_{n} \mid \psi_{n}\right\rangle\right\rangle-M_{0} \tag{8.10}
\end{equation*}
$$

and for odd $k$, i.e. $k=2 n+1$, we find:

$$
\begin{equation*}
M_{2 n+1}=2\left\langle\left\langle\psi_{n} \mid \psi_{n+1}\right\rangle\right\rangle-M_{1} \tag{8.11}
\end{equation*}
$$

Clearly we only have to compute $\psi_{n}$ 's for $n=0, \ldots, K / 2$.
9. Algorithm for calculating the trace of $F(\hat{H} ; p)$

The following steps are necessary:
(1) Note that $p$ is a set of parameters, like $\beta, \mu$ for the electron number function $N_{e}$
(2) Prepare a grid representation of wave functions and the Hamiltonian operator
(3) Estimate $E_{\text {min }}$ and $E_{\text {max }}$ (see: section 7).
(4) Define $\hat{H}_{N}$ (Eq. 7.6).
(5) For a given set of parameters $p$ prepare the Chebyshev coefficients $b_{k}(p)$ $k=0, \ldots, K-1 . K$ (see Eq. (6.6)) has to be even and large enough so that all $b_{k}(p)$ with $k>K$ are negligible (e.g. less than $10^{-8}$ ).
(6) Generate the moments $M_{k} k=0, \ldots, K-1$ :
(a) Set
(i) $M_{0}=N_{g}$
(ii) for all $k=1, \ldots, K-1 M_{k}=0$
(b) Generate a stochastic wave function $\psi_{0}$ on the grid (Eq. 8.2). Note it must be zero on the grid boundaries.
(c) Set
(i) $\psi_{1}=\hat{H}_{N} \psi_{0}$
(ii) $M_{1}=\left\langle\psi_{0} \mid \psi_{1}\right\rangle$,
(iii) $M_{2}=2\left\langle\left\langle\psi_{1} \mid \psi_{1}\right\rangle\right\rangle-M_{0}$;
(d) For $n=2,3, \ldots$ to $n=K / 2$ (remember $K$ is even):
(i) $\psi_{2}=2 \hat{H}_{N} \psi_{1}-\psi_{0}$
(ii) $M_{2 n-1}=M_{2 n-1}+\left(2\left\langle\psi_{n-1} \mid \psi_{n}\right\rangle-M_{1}\right)$
(iii) $M_{2 n}=M_{2 n}+\left(2\left\langle\psi_{n} \mid \psi_{n}\right\rangle-M_{0}\right)$,
(iv) $\psi_{0}=\psi_{1}, \psi_{1}=\psi_{2}$
(e) Average $M_{k}$ over $I$ stochastic functions (see Eq. 8.2 on page 7 ). The statistical error should decrease in proportion to $I^{-1 / 2}$.
(7) Calculate the trace as: $\operatorname{tr}[F(\hat{H} ; p)]=\sum_{k=0}^{K-1} b_{k}(p) M_{k}$.

## References

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