

Stochastic Self-Consistent Second-Order Green's Function Method for Correlation Energies of Large Electronic Systems

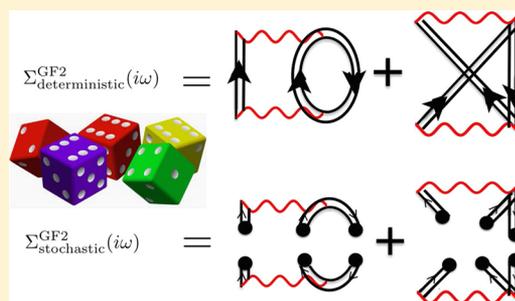
Daniel Neuhauser,^{*,†} Roi Baer,^{*,‡} and Dominika Zgid^{*,§}

[†]Department of Chemistry and Biochemistry, University of California at Los Angeles, Los Angeles, California 90095, United States

[‡]Fritz Haber Center for Molecular Dynamics, Institute of Chemistry, The Hebrew University of Jerusalem, Jerusalem 91904, Israel

[§]Department of Chemistry, University of Michigan, Ann Arbor, Michigan 48109, United States

ABSTRACT: The second-order Matsubara Green's function method (GF2) is a robust temperature-dependent quantum chemistry approach, extending beyond the random-phase approximation. However, until now the scope of GF2 applications was quite limited as they require computer resources that rise steeply with system size. In each step of the self-consistent GF2 calculation there are two parts: estimating of the self-energy from the previous step's Green's function and updating the Green's function from the self-energy. The first part formally scales as the fifth power of the system size, while the second has a much gentler cubic scaling. Here, we develop a stochastic approach to GF2 (sGF2), which reduces the fifth power scaling of the first step to merely quadratic, leaving the overall sGF2 scaling as cubic. We apply the method to linear hydrogen chains with up to 1000 electrons, showing that the approach is numerically stable, efficient, and accurate. The stochastic errors are very small, on the order of 0.1% or less of the correlation energy for large systems, with only a moderate computational effort. The first iteration of GF2 is an MP2 calculation that is done in *linear scaling*; hence we obtain an extremely fast stochastic MP2 (sMP2) method as a byproduct. While here we consider finite systems with large band gaps where at low temperatures effects are negligible, the sGF2 formalism is temperature dependent and general and can be applied to finite or periodic systems with small gaps at finite temperatures.



1. INTRODUCTION

Second-order Green's function (GF2) is a temperature-dependent self-consistent perturbation approach where the Green's function is iteratively renormalized. At self-consistency, the self-energy, which accounts for the many-body correlation effects, is a functional of the Green's function, $\Sigma(G)$. The GF2 approximation as implemented here is described by the diagrams in Figure 1 and employs Matsubara Green's functions that are temperature dependent and expressed on the imaginary axis.^{1–4} The implementation we discuss, for total energies, relies on thermal Matsubara Green's functions instead of real

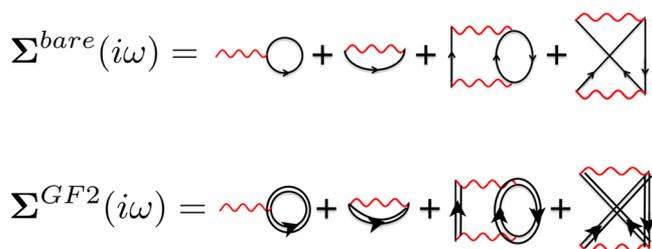


Figure 1. Upper panel: Bare second-order self-energy diagrams. Lower panel: second-order self-energy diagrams evaluated self-consistently. Note that the Green's function lines are renormalized while the interactions lines remain bare. For details, see ref 3.

time Green's functions.^{5–7} This offers advantages in terms of stability and smoothness of the self-energy.

Upon convergence, the GF2 method includes all second-order skeleton diagrams dressed with the renormalized second-order Green's function propagators, as illustrated in Figure 1. Specifically, as shown in ref 3, GF2, which at convergence is reference independent, preserves the desirable features of Møller–Plesset perturbation theory (MP2) while avoiding the divergences that appear when static correlation is important. Additionally, GF2 possesses only a very small fractional charge and spin error,⁴ less than either typical hybrid density functionals or RPA with exchange, therefore having a minimal many-body self-interaction error. In solids, GF2 describes the insulating and Mott regimes and recovers the internal and free energy for multiple solid phases.^{8,9} Moreover, GF2 is useful for efficient Green's function embedding techniques such as in the self-energy embedding method (SEET).^{10–14}

The formal advantages of GF2 come, however, with a price tag. The calculation of the self-energy matrix scales as $O(n_\tau N^5)$, where n_τ is the size of the imaginary time grid and N is the number of atomic orbitals (AOs). This leads to steep numerical costs, which prevent practical application of GF2 to systems larger than a few dozen electrons. The practical application to larger systems requires therefore a different paradigm, and here

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we therefore develop a statistical formulation of GF2 that calculates the self-energy matrix in linear-scaling.

The key to the present development, distinguishing it from previous work,^{15–17} is the conversion of nested summations into stochastic averages. Our method draws from previous work on stochastic electronic structure methods, including stochastic density functional theory (sDFT),^{18–20} sDFT with long-range exact exchange,²¹ multiexciton generation,²² Møller–Plesset perturbation theory (sMP2),^{23–25} random-phase approximation (sRPA),²⁶ GW approximation (sGW),^{27–29} time-dependent DFT (sTDDFT),³⁰ optimally tuned range separated hybrid DFT,³¹ and Bethe–Salpeter equation (sBSE).³² Among these, the closest to this work are the stochastic version of sMP2 in real-time plane-waves^{23,24} and MO-based MP2 with Gaussian basis sets.²⁴ The stochastic method presented here benefits from the fact that the GF2 self-energy is a smooth function of imaginary time and is therefore naturally amenable to random sampling.

2. METHOD

2.1. Brief Review of GF2. Our starting point is a basis of N real single-electron nonorthogonal atomic-orbital (AO) states, $\phi_i(\mathbf{r})$, with an $N \times N$ overlap matrix $S_{ij} = \langle \phi_i | \phi_j \rangle$. Such states could be of any form, Gaussian, numerical, etc., but for efficiency should be localized. We then use second quantization creation, a_i^\dagger , and annihilation, a_i , operators with respect to the nonorthogonal basis $\phi_i(\mathbf{r})$. The nonorthogonality is manifested only in a modified commutation relation,

$$\{a_i, a_j^\dagger\} = (S^{-1})_{ij} \quad (1)$$

The Hamiltonian for the interacting electrons has the usual form

$$\hat{H} = \sum_{ij} h_{ij} a_i^\dagger a_j + \frac{1}{2} \sum_{ijkl} v_{ijkl} a_i^\dagger a_k^\dagger a_l a_j \quad (2)$$

where $h_{ij} = \int d\mathbf{r} \phi_j(\mathbf{r})(-1/2\nabla^2 + v_{\text{ext}}(\mathbf{r}))\phi_i(\mathbf{r})$ and $v_{\text{ext}}(\mathbf{r})$ is the bare external potential (due to the nuclei), while \hat{V} is the two electron–electron (e–e) Coulomb interaction described by the 2-electron integrals

$$v_{ijkl} = \int \int \phi_i(\mathbf{r})\phi_j(\mathbf{r})v(\mathbf{r}-\mathbf{r}')\phi_k(\mathbf{r}')\phi_l(\mathbf{r}') d\mathbf{r} d\mathbf{r}' \quad (3)$$

where $v(\mathbf{r}) = \frac{1}{r}$ is the Coulomb interaction potential.

At a finite temperature β^{-1} and chemical potential μ , we employ the grand canonical density operator $\frac{e^{-\beta(\hat{H}-\mu\hat{N})}}{Z}$, where $\hat{N} = \sum_{ij} S_{ij} a_i^\dagger a_j$ is the electron-number operator and $Z(\beta) = \text{Tr}[e^{-\beta(\hat{H}-\mu\hat{N})}]$ is the partition function. The thermal expectation value of any operator \hat{A} can be calculated as $\langle \hat{A} \rangle = Z^{-1}(\beta) \text{Tr}[e^{-\beta(\hat{H}-\mu\hat{N})} \hat{A}]$. For one-body observables $\hat{A} = \sum_{ij} A_{ij} a_i^\dagger a_j$, we write $\langle \hat{A} \rangle = \sum_{ij} A_{ij} P_{ij}$ where $P_{ij} = \langle a_i^\dagger a_j \rangle$ is the reduced density matrix.

The 1-particle Green's function $G_{jk}(\tau)$ at an imaginary time τ is a generalization of the concept of the density matrix and obeys an equation of motion that can be solved by perturbation methods. Formally:

$$G_{jk}(\tau) = -\langle T a_j(\tau) a_k^\dagger \rangle \quad (4)$$

where $a_j(\tau) \equiv e^{(\hat{H}-\mu\hat{N})\tau} a_j e^{-(\hat{H}-\mu\hat{N})\tau}$ with $-\beta < \tau < \beta$, and T is the time-ordering symbol:

$$T a_j(\tau) a_k^\dagger \equiv \theta(\tau) a_j(\tau) a_k^\dagger - \theta(-\tau) a_k^\dagger a_j(\tau) \quad (5)$$

Note that $\mathbf{G}(\tau)$ is a real symmetric matrix.

Each element $G_{jk}(\tau)$ (and therefore the entire matrix $\mathbf{G}(\tau)$) is discontinuous when going from negative to positive times, but this discontinuity is not a problem since we only need to treat explicitly positive times $\tau > 0$, while negative τ 's are accessible by the antiperiodic relation for $\mathbf{G}(\tau)$

$$\mathbf{G}(\tau) = -\mathbf{G}(\tau + \beta), \quad -\beta < \tau < 0 \quad (6)$$

as directly verified by substitution in eq 4. Hence $\mathbf{G}(\tau)$ can be expanded as a Fourier series involving the Matsubara frequencies $\omega_n = (2n + 1)\frac{\pi}{\beta}$:

$$\mathbf{G}(\tau) = \frac{1}{\beta} \sum_{n=-\infty}^{\infty} \mathbf{G}(i\omega_n) e^{-i\omega_n\tau} = \frac{2}{\beta} \text{Re} \sum_{n=0}^{\infty} \mathbf{G}(i\omega_n) \quad (7)$$

where:

$$\mathbf{G}(i\omega_n) = \int_0^\beta \mathbf{G}(\tau) e^{i\omega_n\tau} d\tau \quad (8)$$

where the second equality in eq 7 is due to the fact that $\mathbf{G}(\tau)$ is real.

The Green's function of eq 4 gives access to the reduced density matrix by taking the imaginary time τ as a negative infinitesimal (denoted as 0^-):

$$P_{kj} = 2G_{kj}(0^-) = -2G_{kj}(\beta^-) = \frac{4}{\beta} \text{Re} \sum_{n=0}^{\infty} e^{-i\omega_n 0^-} G_{kj}(i\omega_n) \quad (9)$$

Hence, all thermal averages of one-electron operators are accessible through the sum of the Matsubara coefficients.

Perturbation theory can be used to build approximations for $\mathbf{G}(\tau)$ based on a noninteracting Green's function $\mathbf{G}_0(\tau)$ corresponding to a reference one-body Hamiltonian $\hat{H}_0 = \sum_{ij} F_{ij} a_i^\dagger a_j$. Here, \mathbf{F} is any real symmetric "Fock" matrix such that \hat{H}_0 well approximates the interacting electron Hamiltonian. The derivation of $\mathbf{G}_0(\tau)$ requires orthogonal combination of the basis set, that is, finding a matrix \mathbf{X} that fulfills $\mathbf{X}\mathbf{X}^T = \mathbf{S}^{-1}$. Then it is straightforward to show that

$$\mathbf{G}_0(\tau) = \mathbf{X} e^{-\tau(\bar{\mathbf{F}}-\mu\mathbf{I})} \left[\frac{\theta(-\tau)}{1 + e^{\beta(\bar{\mathbf{F}}-\mu\mathbf{I})}} - \frac{\theta(\tau)}{1 + e^{-\beta(\bar{\mathbf{F}}-\mu\mathbf{I})}} \right] \mathbf{X}^T \quad (10)$$

where $\bar{\mathbf{F}} = \mathbf{X}^T \mathbf{F} \mathbf{X}$ is the Fock matrix in the orthogonal basis set. Note that for positive (or negative) imaginary times $\mathbf{G}_0(\tau)$ is a real, smooth and nonoscillatory Green's function. This is important for us since it is much easier to stochastically sample a smooth function.

Performing the integral of eq 8 with $\mathbf{G}_0(\tau)$ of eq 10 yields

$$\mathbf{G}_0(i\omega_n) = ((\mu + i\omega_n)\mathbf{S} - \mathbf{F})^{-1} \quad (11)$$

Since we now know how to write down Green's functions for noninteracting systems, we rewrite the unknown part of the exact Green's function by introducing the frequency-dependent self-energy, formally defined by

$$\mathbf{G}(i\omega_n) = ((\mu + i\omega_n)\mathbf{S} - \mathbf{F} - \boldsymbol{\Sigma}(i\omega_n))^{-1} \quad (12)$$

and by construction the self-energy fulfills the Dyson equation:

$$\mathbf{G}(i\omega_n) = \mathbf{G}_0(i\omega_n) + \mathbf{G}_0(i\omega_n)\boldsymbol{\Sigma}(i\omega_n)\mathbf{G}(i\omega_n) \quad (13)$$

Instead of viewing these equations as a definition of the self-energy $\Sigma(i\omega_n)$, we can calculate this self-energy to a given order of perturbation theory in $\Delta\hat{H} = \hat{H} - \hat{H}_0$. Specifically, the GF2 approximation^{2,3} uses a Hartree–Fock ansatz for \mathbf{F} ,

$$F_{ij} = h_{ij} + \frac{1}{2}P_{kl}(2v_{ijkl} - v_{ilkj}) \quad (14)$$

where an Einstein summation convention is used, summing indices that appear in pairs (here, both k and l). The self-energy in imaginary time $\Sigma(\tau)$ is then obtained by second-order perturbation theory (see Figure 1):

$$\Sigma_{ij}(\tau) = G_{kl}(\tau)G_{mn}(\tau)G_{pq}(\beta - \tau)v_{impk}(2v_{mlq} - v_{lnq}) \quad (15)$$

Note that $\Sigma(\tau)$ and $\Sigma(i\omega_n)$ are connected by exactly the same Matsubara relations connecting $\mathbf{G}(\tau)$ and $\mathbf{G}(i\omega_n)$, eqs 7–8.

The self-consistent one-body Green's function governs all one-body expectation values. Moreover, even the total two-body potential energy is available, by differentiation of the matrix trace (denoted by $\text{Tr}[\cdot]$) of the Green's function with respect to τ : $\langle \hat{V} \rangle = -\frac{1}{2}\lim_{\tau \rightarrow 0^-} \text{Tr}\left[\left(\mathbf{S}\left(\frac{\partial}{\partial \tau} - \mu\right) + \mathbf{h}\right)\mathbf{G}(\tau)\right]$.

Hence, the total energy is

$$\langle \hat{H} \rangle = \text{Tr}\left[\mathbf{h}\mathbf{P} - \frac{1}{2}\lim_{\tau \rightarrow 0^-}\left[\left(\frac{\partial}{\partial \tau} - \mu\right)\mathbf{S} + \mathbf{h}\right]\mathbf{G}(\tau)\right] \quad (16)$$

Plugging the definition of $\mathbf{G}(\tau)$ into to eq 16 the total energy has simple time and frequency forms:

$$\begin{aligned} \langle \hat{H} \rangle &= \frac{1}{2}\text{Tr}[(\mathbf{h} + \mathbf{F})\mathbf{P}] + \frac{2}{\beta}\text{Re} \sum_{n \geq 0} \text{Tr}[\mathbf{G}(i\omega_n)\Sigma^T(i\omega_n)] \\ &= \frac{1}{2}\text{Tr}[(\mathbf{h} + \mathbf{F})\mathbf{P}] + \int_0^\beta \text{Tr}[\mathbf{G}(\beta - \tau)\Sigma(\tau)] d\tau \end{aligned} \quad (17)$$

To conclude, the combination of eqs 9, 12, 14, and 15 along with the requirement that the density matrix describes N_e electrons results in the following self-consistent GF2 procedure:

1. Perform a standard HF calculation and obtain a starting guess for the Fock matrix $\mathbf{F} = \mathbf{F}_{\text{HF}}$ and the density matrix $\mathbf{P} = \mathbf{P}_{\text{HF}}$. Set $\Sigma(i\omega_n) = 0$ for the set of N_ω positive Matsubara frequencies ω_n , $n = 0, 1, 2, \dots, N_\omega - 1$.
2. Given $\Sigma(i\omega_n)$ and \mathbf{F} , find μ such that $\text{Tr}[\mathbf{P}\mathbf{S}] = N_e$ where \mathbf{P} is given in eq 9 from $\mathbf{G}(\tau = \beta^-)$, which depends on μ through the basic definition eq 12.
3. Calculate $\mathbf{G}(\tau)$ (eq 7) and \mathbf{P} (eq 9).
4. Calculate the Fock matrix \mathbf{F} from \mathbf{P} (eq 14).
5. Calculate the self-energy $\Sigma(\tau)$ from eq 15 and transform is to the Matsubara frequency domain to yield $\Sigma(i\omega_n)$.
6. Calculate the total energy $\langle \hat{H} \rangle$ from eq 17.
7. Repeat steps 2–6 until convergence of the density and the total energy.

Once converged, the GF2 correlation energy is defined as the difference $E_{\text{corr}} = \langle \hat{H} \rangle - E_{\text{HF}}$ between the converged total energy (eq 17) and the initial Hartree–Fock energy, $E_{\text{HF}} = \frac{1}{2}\text{Tr}[(\mathbf{h} + \mathbf{F}_{\text{HF}})\mathbf{P}_{\text{HF}}]$. Note that in the first iteration GF2 yields automatically the temperature-dependent MP2 energy:

$$E_{\text{MP2}}^{\text{corr}} = \int_0^\beta \text{Tr}[\mathbf{G}_0(\beta - \tau; \mathbf{F}_{\text{HF}})\Sigma_0(\tau)] d\tau \quad (18)$$

where $\Sigma_0(\tau)$ is that of eq 15 with \mathbf{G}_0 replacing \mathbf{G} . This expression reduces to the familiar MP2 energy expression at the limit $\beta \rightarrow \infty$ (zero temperature limit), when evaluated in the molecular orbital basis set that diagonalizes the matrix \mathbf{F}_{HF} .

Finally, a technical point. The representation of the Green's functions in τ -space can be complicated when the energy range of the eigenvalues of \mathbf{F} is large since a function of the type $e^{-\tau(f-\mu)}/(1 + e^{-\beta(f-\mu)})$ can be spiky when $f > \mu$ and $\tau \rightarrow \beta$ or when $f < \mu$ and $\tau \rightarrow 0$. This requires special techniques for both imaginary time and frequency grids as discussed in refs 33 and 34.

2.2. sGF2: Stochastic Approach to GF2. Most of the computational steps in the above algorithm scale with system size N (number of AO basis functions) as $O(N_{\text{SC}} \times N_\tau \times N^3)$ where N_{SC} is the number of GF2 self-consistent iterations and N_τ is the number of time-steps. However, the main numerical challenge in GF2 is step 5 (eq 15), which scales formally as $O(N_{\text{SC}} \times N_\tau \times N^5)$ making GF2 highly expensive for any reasonably sized system. This steep scaling is due to the contraction of two 4-index tensors with three Green's function matrices.

To reduce this high complexity, we turn to the stochastic paradigm, which represents the matrices $\mathbf{G}(\tau)$ by an equivalent random average over stochastically chosen vectors. Fundamentally, this is based on resolving the identity operator. Specifically, for each τ we generate a vector $\boldsymbol{\eta}^0$ of N components randomly set to +1 or -1. Vectors at different times τ are statistically independent, but we omit for simplicity their τ labeling. Then, the key, and trivial, observation is that average of the product of different components of $\boldsymbol{\eta}^0$ is the unit matrix, which we write symbolically as

$$\eta_k^0 \eta_l^0 = \delta_{kl} \quad (19)$$

We emphasize that the equality in this equation should be interpreted to hold in the limit of averaging over infinitely many random vectors $\boldsymbol{\eta}^0$.

Given this separable presentation of the unit matrix, it is easy to rewrite any matrix as an average over separable vectors. Specifically, from $\boldsymbol{\eta}^0$ we define the two vectors:

$$\boldsymbol{\eta} = \sqrt{|\mathbf{G}(\tau)|}\boldsymbol{\eta}^0, \quad \bar{\boldsymbol{\eta}} = \text{sgn}(\mathbf{G}(\tau))\sqrt{|\mathbf{G}(\tau)|}\boldsymbol{\eta}^0 \quad (20)$$

and then

$$G_{kl}(\tau) = \bar{\eta}_k \eta_l \quad (21)$$

Here, the square-root matrix is $\sqrt{|\mathbf{G}(\tau)|} = \mathbf{A}\sqrt{|\mathbf{g}|}\mathbf{A}^T$, where $\mathbf{A}(\tau)$ is the unitary matrix of eigenvectors and $\mathbf{G}(\tau)$ is the diagonal matrix of eigenvalues of $\mathbf{G}(\tau)$.

As a side note, we have a freedom to choose other vectors; specifically, any two vectors $\bar{\boldsymbol{\eta}} = \bar{\mathbf{D}}\boldsymbol{\eta}^0$, $\boldsymbol{\eta} = \mathbf{D}\boldsymbol{\eta}^0$, will work if $\bar{\mathbf{D}}\mathbf{D}^T = \mathbf{G}(\tau)$. In principle, we can even use the simplest choice $\bar{\mathbf{D}} = 1$, $\mathbf{D} = \mathbf{G}(\tau)$, corresponding to $\bar{\boldsymbol{\eta}} = \boldsymbol{\eta}^0$ and $\boldsymbol{\eta} = \mathbf{G}(\tau)\boldsymbol{\eta}^0$. But while this latter choice has the advantage that $\mathbf{G}(\tau)$ does not need to be diagonalized, we find that it is numerically better to use eq 20 as it is more balanced and therefore converges faster with the number of stochastic samples. Also note that at the first iteration, where $\mathbf{G}(\tau) = \mathbf{G}_0(\tau)$, there is no need to diagonalize $\mathbf{G}_0(\tau)$ at different times, since it is obtained directly from the eigenstates of \mathbf{F} in eq 10.

Going back to eq 20, we similarly separate the other two Green's function matrices appearing in eq 15, writing them as $G_{mn}(\tau) = \bar{\xi}_m \xi_n$ and $G_{pq}(\beta - \tau) = \bar{\zeta}_p \zeta_q$. The self-energy in eq 15 is then

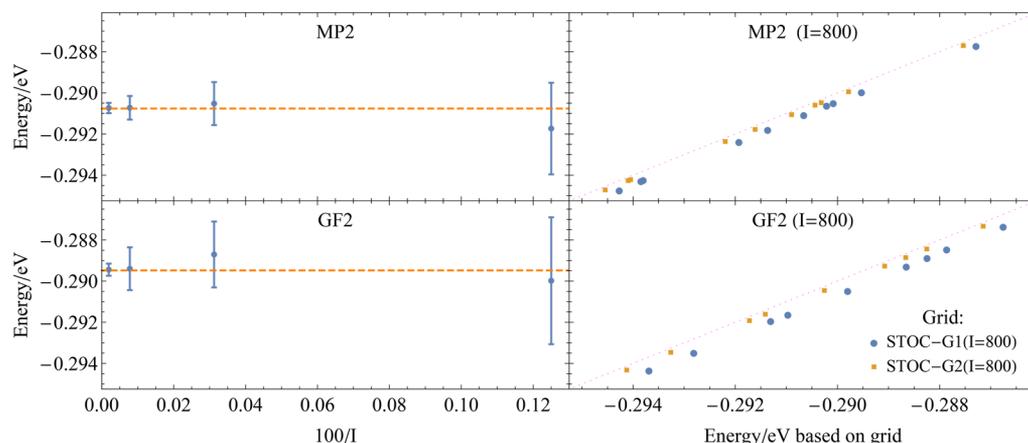


Figure 2. MP2 and GF2 correlation energies per electron for a linear H_{10} system with nearest neighbor spacing of 1 Å. Left panels: the DET (deterministic) correlation energies (dashed horizontal lines) are well within the I -dependent error-bars $\bar{E} \pm 2\sigma$ of the STOC-NG calculations, where \bar{E} and σ are the average and standard deviation of the correlation energy calculated in 10 statistically independent runs. Right panels: a correlation plot of pairs of stochastically ($I = 800$) estimated correlation energies: ($E_{\text{STOC-G1}(800)}$, $E_{\text{STOC-NG}(800)}$); as blue dots and ($E_{\text{STOC-G2}(800)}$, $E_{\text{STOC-NG}(800)}$); as orange dots, $i = 1, \dots, 10$. The diagonal dotted line represents the perfect correlation $E_{\text{STOC-G1-or-2}} = E_{\text{STOC-NG}}$.

$$\Sigma_{ij}(\tau) = \bar{\eta}_k \bar{\xi}_m \bar{\zeta}_p v_{impk} (2\eta_l \xi_n \zeta_q v_{jnlq} - \eta_l \xi_n \zeta_q v_{jlnq}) \quad (22)$$

so that it is separable to a product of two terms

$$\Sigma_{ij}(\tau) = \bar{u}_i [2u_j - w_j] \quad (23)$$

where we defined three auxiliary vectors

$$\begin{aligned} \bar{u}_i &= \bar{\eta}_k \bar{\xi}_m \bar{\zeta}_p v_{impk} \\ u_j &= \eta_l \xi_n \zeta_q v_{jnlq} \\ w_j &= \eta_l \xi_n \zeta_q v_{jlnq} \end{aligned} \quad (24)$$

The self-energy in eq 23 should be viewed as the average, over the stochastic vectors ξ^0 , η^0 , and ζ^0 , of the product term (\bar{u}_i times $2u_j - w_j$).

The direct calculation of the vectors \bar{u} , u , w by eq 24 is numerically expensive once $M > 30$. We reduce the scaling by recalling the definition of v_{jnlq} in eq 3:

$$u_j = \eta_l \xi_n \zeta_q \iint \phi_j(\mathbf{r}) \phi_n(\mathbf{r}) v(|\mathbf{r} - \mathbf{r}'|) \phi_l(\mathbf{r}') \phi_q(\mathbf{r}') \, d\mathbf{r} \, d\mathbf{r}' \quad (25)$$

$$= \iint \phi_j(\mathbf{r}) \xi(\mathbf{r}) v(|\mathbf{r} - \mathbf{r}'|) \eta(\mathbf{r}') \zeta(\mathbf{r}') \, d\mathbf{r} \, d\mathbf{r}' \quad (26)$$

where

$$\eta(\mathbf{r}) = \eta_l \phi_l(\mathbf{r}) \quad (27)$$

and $\xi(\mathbf{r})$ and $\zeta(\mathbf{r})$ are analogously defined. We can therefore write

$$u_j = \int \phi_j(\mathbf{r}) \xi(\mathbf{r}) v_{\eta\zeta}(\mathbf{r}) \, d\mathbf{r} \quad (28)$$

where

$$v_{\eta\zeta}(\mathbf{r}) \equiv \int v(|\mathbf{r} - \mathbf{r}'|) \eta(\mathbf{r}') \zeta(\mathbf{r}') \, d\mathbf{r}' \quad (29)$$

is the Coulomb potential corresponding to the random charge distribution $\eta(\mathbf{r})\zeta(\mathbf{r})$. Similar expressions apply for \bar{u}_i and w_j .

Equations 28 and 29 are performed numerically using FFT methods on a 3D Cartesian grid with N_g grid points, so eq 29 is calculated with $O(N_g \log N_g)$ operations. Since the AO basis

functions $\phi_i(\mathbf{r})$ are local in 3D space, the calculations of $\eta(\mathbf{r})$, $\xi(\mathbf{r})$, and $\zeta(\mathbf{r})$ in eq 27 scale linearly with system size.

Equation 23 gives an exact expression for $\Sigma_{ij}(\tau)$, as an expected value over formally an infinite number of stochastic orbitals η^0 , ξ^0 , and ζ^0 . Actual calculations use a finite number I of “stochastic iterations”, where in each such iteration a set of stochastic vectors η^0 , ξ^0 , and ζ^0 (different at each τ) is generated and $\Sigma_{ij}(\tau)$ is averaged over them. The overall scaling of this step is therefore $I \times N_r \times (N_g \log N_g + N^2)$. We note that the typical values of N_r and I are in the hundreds, see the discussion of the stochastic error below.

Finally, we note that while the stochastic vectors (η^0 , ξ^0 , ζ^0) are statistically independent for each time point τ , the same τ -dependent vectors are used at each GF2 iteration, making it possible to converge these iterations.

3. RESULTS

3.1. Systems and Specifics. The algorithm was tested on linear hydrogen chains (H_M) with a nearest neighbor distance of 1 Å, for several sizes: $M = 10, 100, 300$, and 1000. The linearity was for convenience, and we emphasize that it does not play any role in the algorithm. The smallest chain was used to demonstrate the convergence of the approach to the basis-set deterministic values, and the other three calculations were used to study the dependence of the algorithm on system size.

In all calculations, an STO-3G basis was used, so that in this case $N = M$ and obviously the number of electrons is also $N_e = M$. A periodic spatial grid of $0.5a_0$ spacing was used to represent the wave functions, and the grids contained 10×10 points in the direction orthogonal to chain and between 60 and 4000 points along the chain, depending on system size. For the smallest system (H_{10}) a finer, bigger grid was also used, as detailed below.

Other, technical details:

- Periodic images were screened using the method of ref 35.
- The inverse temperature was $\beta = 50E_h^{-1}$.
- A Chebyshev-type imaginary-time grid with 128 time points was employed using a spline-fit method^{33,34} for the frequency-to-time conversions of $G(i\omega_n)$ and $\Sigma(i\omega_n)$ and for the evaluation of the two-body energy.

3.2. Small System. In our GF2 and MP2 algorithm, we make two types of numerical discretizations. First, we use a finite number (labeled I) of stochastic iterations to sample the self-energy, so we must show convergence as I grows. Second, we use grids for bypassing the need to sum over $O(N^4)$ two-electron integrals; hence we need to demonstrate convergence with respect to grid quality. We therefore examine in this section a small system, linear H_{10} , and make four types of GF2/MP2 correlation energy calculations:

- DET, fully deterministic calculations based on the analytical 2-electron integrals;
- STOC(I)-NG, stochastic calculations based on I stochastic iterations and on the analytical two electron integrals, and NG stands for No Grid;
- STOC(I)-G1 and STOC(I)-G2, stochastic calculations based on I stochastic iterations and on a 3D grid. Here, G1 is the same type of grid we use for the larger calculations, and includes $10 \times 10 \times 60$ points with a spacing $h = 0.5a_0$. G2 is somewhat denser and covers more space, with $16 \times 16 \times 100$ points and $h = 0.4a_0$.

Our strategy is to first show that STOC-NG(I) converges to the deterministic set (DET) as I grows. Then we show that for a given number of stochastic orbitals, $I = 800$, both grid results are quite close to the nongrid result and that the somewhat better second grid (STOC($I = 800$)-G2 leads to extremely close results to the nongrid values (STOC($I = 800$)-NG), so that the convergence with grid is very rapid.

We repeat the STOC-NG calculation 10 times determining the average correlation energy \bar{E} and its standard deviation σ as a function of I . The results are shown in the left panels of Figure 2 as error-bars at $\bar{E} \pm \sigma$, which shrink approximately as $1/\sqrt{I}$ and which include the DET result, represented as dashed horizontal lines, showing very small or no bias. For MP2, a bias in the stochastic calculations is not expected since the correlation energy is calculated linearly from the first iteration of the self-energy Σ_0 (eq 18). But for GF2 such a bias may form since the “noisy” self-energy is used nonlinearly to update the Green’s function in eq 12. However, for this small $N = 10$ system the stochastic MP2 and GF2 energies do not exhibit a noticeable bias. We discuss the bias in larger systems below.

Next, we assess the errors associated with using grid calculations replacing the analytical 2-electron integration. In both right panels of Figure 2, we show 10 blue dots, each corresponding to a pair of stochastic energies $(E_{\text{STOC-G1}(800)}, E_{\text{STOC-NG}(800)})_i$, $i = 1, \dots, 10$, both calculated with the same random seed s_i (of course, s_i and s_j are statistically independent). We also show 10 orange dots, each corresponding to a pair of stochastic energies $(E_{\text{STOC-G2}(800)}, E_{\text{STOC-NG}(800)})_i$, also calculated with the same seed s_i as before. The use of the same seeds for each pair of blue and orange dots allows for comparison of the grid error (which is the horizontal distance of a point from the diagonal) without worrying about the larger statistical error, seen as the spread of the results along the diagonal. We see that the grid error decreases significantly when moving from G1 to G2, but even the error for G1 is already very small (about 0.5 meV per electron).

3.3. Larger Systems. In the small system considered above, the bias was not noticeable, and here we examine the bias in larger systems. In Figure 3, we show the STOC-G1(I) correlation energies in three specific systems composed of $N = 100, 300,$ and 1000 hydrogen atoms placed on a straight line with a nearest neighbor spacing of 1 \AA .

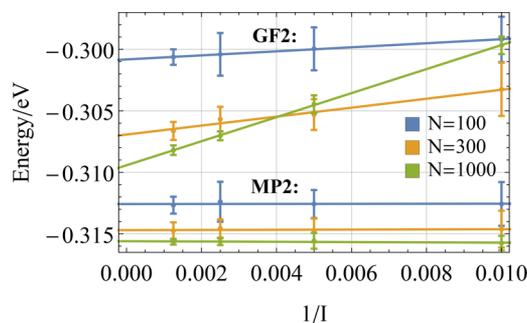


Figure 3. Statistical estimates of the MP2 and GF2 correlation energies per electron for three linear chains with $N = 100, 300,$ and 1000 H atoms as described in the text, as a function of the inverse number of sets of stochastic orbitals I . The results are shown as error bars, where the center of each error bar is the average and its width is the standard deviation of the correlation energy estimates from 10 statistically independent runs, each employing I sets of stochastic orbitals. The lines are linear regression fits to the data and uncertainties.

We first study the MP2 correlation energy of each system, appearing in the lower energy range in the figure. The starting point of the GF2 calculation is the Hartree–Fock F_{HF} and P_{HF} matrices, so the MP2 energy is half the correlation energy of the first self-consistent iteration (see eqs 17 and 18). The statistical errors in MP2 are pure fluctuations, a random number distributed normally with zero average and with standard deviation given by σ_0/\sqrt{I} where σ_0 is independent of I but shrinks with chain length: $\sigma_0 \propto 1/\sqrt{L}$, exhibiting “self averaging”.¹⁸ The stochastic MP2 errors are very small and decrease with system size, so for $N = 1000$ the standard deviation of the $I = 800$ iteration calculation is 0.07% of the total correlation energy. For perspective, note that (deterministic) errors of larger or similar magnitude are present in linear scaling local or divide and conquer MP2 methods with density fitting.^{36,37}

Next, we discuss the stochastic estimates of the self-consistent GF2 correlation energies. These exhibit statistical errors with two visible components. The first is a *fluctuation*, similar in nature to that of the MP2 calculation, and the second component is a *bias*, which decreases as I grows. In fact, we expect the bias to asymptotically decrease inversely with I ,^a so we fit the numerical GF2 results to a straight line in I^{-1} . Table 1 shows the estimate of the correlation energies, the fluctuation, and the bias as a function of the number of stochastic orbitals. The results are highly accurate; for example, when $I = 800$ is used for the largest system ($N = 1000$), the errors in MP2 and in GF2 are smaller than 0.1%.

3.3.1. Timings. The measured overall CPU time for the stochastic self-energy calculation (performed on a XEON system) can be expressed as

$$T^\Sigma \approx 2.5 \times N \times N_\tau \times I \times 10^{-7} \text{ h} \quad (30)$$

where, as mentioned, N is the number of electrons and I the number of stochastic orbitals in the system. The MP2 wall time calculation is essentially equal to the self-energy time divided by the number of cores, n_{CORES} , since the parallelization has negligible overhead:

$$T_{\text{wall}}^{\text{MP2}} \approx \frac{T^\Sigma}{n_{\text{CORES}}} \quad (31)$$

Table 1. Statistical Estimate of the MP2 and GF2 Energies Per Electron of the Linear Hydrogen Chains H_N , Based on Runs with up to $I = 800$ Stochastic Samplings^a

| N | MP2 energy (eV) | | GF2 energy (eV) | |
|------|-----------------------------|--------------|--|--------------|
| | $E(N,I)$ | $I = 800$ | $E(N,I)$ | $I = 800$ |
| 100 | $-0.3126 \pm 0.025I^{-1/2}$ | $-0.3127(9)$ | $-0.3008 + 0.16I^{-1} \pm 0.025I^{-1/2}$ | $-0.3007(9)$ |
| 300 | $-0.3148 \pm 0.015I^{-1/2}$ | $-0.3148(5)$ | $-0.3069 + 0.36I^{-1} \pm 0.020I^{-1/2}$ | $-0.3067(7)$ |
| 1000 | $-0.3157 \pm 0.007I^{-1/2}$ | $-0.3157(2)$ | $-0.3094 + 0.98I^{-1} \pm 0.010I^{-1/2}$ | $-0.3094(3)$ |

^aFor MP2, there is no bias and the error is a only statistical fluctuation, which for $I = 800$ is very small, on the order of 0.1% of the correlation energy or better. For GF2, the statistical fluctuation is similarly tiny, and there is a bias (depending to leading order on I^{-1}) that grows with N , and therefore I should be similar to N or larger.

GF2 involves an additional step, where the Green's function is constructed from the self-energy, and this step scales cubically with system size. Furthermore, there are N_{SC} self-consistent iterations. The total time is therefore found to be

$$T_{\text{wall}}^{\text{GF2}} \approx N_{\text{SC}} \frac{(1.7 \times N^3 \times N_r \times 10^{-11} \text{ h} + T^{\Sigma})}{n_{\text{CORES}}}$$

For the H_{1000} system, with $I = 800$ stochastic orbitals, the MP2 calculation takes $T_{\text{wall}}^{\text{MP2}} = 24 \text{ h}/n_{\text{CORES}}$, that is, about 30 min when using 48 cores.

The GF2 calculation for this same system involves $N_{SC} = 12$ iterations and a cubic part which takes about 2 core-hours per iteration, that is, the cubic part is still an order of magnitude smaller than the self-energy sampling time for this system size. The wall time is therefore $T_{\text{wall}}^{\text{GF2}} = 6.5 \text{ h}$ with 48 cores.

For the H_{300} system, we find $T_{\text{wall}}^{\text{MP2}} = 12 \text{ min}$ and $T_{\text{wall}}^{\text{GF2}} = 2 \text{ h}$, while for H_{100} we have $T_{\text{wall}}^{\text{MP2}} = 3.5 \text{ min}$ and $T_{\text{wall}}^{\text{GF2}} = 40 \text{ min}$.

Note that these timings are for a single calculation. The error estimation uses, as mentioned, ten completely independent runs, and therefore took 10 times longer. Of course, once one knows the statistics, it is better to only increase I and not spend CPU time on multiple independent calculations. This is because increasing I reduces the bias while running multiple independent calculations does not. Put differently, in production runs, rather than doing 10 different runs each with a given I , it is better to do a single run with $10I$ stochastic samplings.

For comparison, we note that for H_{100} a deterministic calculation takes 45 min on a single core, which is 4 times faster than the stochastic calculation. Since the deterministic algorithm scales steeply, as $O(N^5)$, the deterministic–stochastic crossover occurs already at H_{150} . Note that for the largest system we studied, H_{1000} , the deterministic calculation would take $\frac{10^4 - 10^5}{n_{\text{CORES}}}$ wall time hours per SCF iteration, compared to $\frac{24}{n_{\text{CORES}}}$ hours for the stochastic calculations.

3.4. Born–Oppenheimer Potential Curves. Potential energy curves can be calculated by correlated sampling, where at each new nuclear configuration one employs the same set of stochastic orbitals η^0 , ξ^0 , and ζ^0 for the self-energy estimation. For demonstration, the HF, MP2, and GF2 Born–Oppenheimer potentials of the H_{100} system are shown in Figure 4 as a function of the displacement of atom no. 25 (counting from the left). In all three methods, the most stable position of the atom is at approximately $-0.1a_0$, slightly displaced toward the nearest chain end. HF theory produces an energy potential with large variations of up to 1.5 eV and large vibrational frequencies of order of 3.4 eV. The MP2 curve is much smoother and the vibrational frequencies reduces to $\sim 2.4 \text{ eV}$, while the GF2

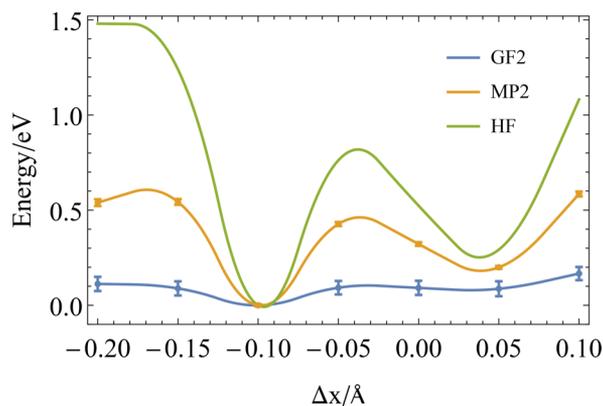


Figure 4. Hartree–Fock, MP2, and GF2 potential energy curves for the energy change in the displacement Δx of atom no. 25 in the 100 H chain.

energy curve is considerably flatter, predicting a vibrational frequency of $\sim 1.0 \text{ eV}$.

4. SUMMARY AND CONCLUSIONS

The problem we addressed here is the reduction of the steep $O(N^5)$ scaling associated with the implementation of self-consistent GF2 calculations. We developed an effective way to reduce the complexity to $O(N^3)$ by using stochastic techniques for calculating the self-energy. A detailed derivation was given along with a specific algorithm. The sampling error in the overall algorithm was studied for linear H_N systems, and the simulation showed that the stochastic errors in the correlation energies can be controlled to less than 0.1% for very large systems. While the studied systems were linear, the algorithm makes no use of the linearity and applies equally well to any geometry.

As a byproduct, since the first step in GF2 is equivalent to MP2, we obtain a stochastic MP2 method (sMP2) performed on top of an existing HF calculation. This approach too has a formal complexity of $O(N^5)$, which is reduced here to *linear* $O(N)$, except for a single overall Fock-matrix diagonalization, which is often available from the underlying HF or DFT ground-state calculation. The errors in this well-scaling stochastic MP2 method are comparable to those of local MP2 approaches used in quantum chemistry.

For GF2, the method has two main stages. The first stage, as in the MP2 case, is a linear scaling calculation of the self-energy. This self-energy is then used in the second stage to construct the Green's function, at an $O(N^3)$ cost. A complication arises in GF2 due to this second stage (but not in MP2!), where the self-energy enters nonlinearly into the expression for the Green's function. This nonlinearity gives rise to a noticeable bias, which is proportional to the system size N . To overcome this bias the

number of stochastic orbitals I used in the first step must be increased in proportion to the system size N , and hence the self-energy calculation in GF2 attains an $O(N^2)$ scaling. The overall scaling of the GF2 calculation is unaffected by this bias problem and remains $O(N^3)$.

The present calculations give a fully self-consistent Green's function method for a large system with a thousand electrons described by a full quantum chemistry Hamiltonian. Moreover, we demonstrated that the splitting of matrices by a random average over stochastically chosen vectors leads to small variance and that relatively few Monte Carlo samples already yield quite accurate correlation energies. The reason for this excellent sampling dependence is 2-fold: the stochastic sampling inherently acts only in the space of atomic orbitals, while the actual spatial integrals (eq 29) are evaluated using a deterministic, numerically exact calculation; in addition, since the Green's function matrices are smooth in imaginary time, different random vectors can be used at each imaginary-time point thereby enhancing the stochastic sampling efficiency.

We have shown that both sMP2 and sGF2 are suitable for calculating potential energy curves or surfaces. Interestingly, for the H_N systems, the potential curve is much smoother and flatter than that in HF or MP2.

As for future applications, we note that sGF2 and sMP2 methods are automatically suitable for periodic systems, as all the deterministic steps and the time-frequency transforms are very efficient when done in the reciprocal (k) space. The only additional detail is that in periodic systems one needs to choose the random vectors to be in k -space and then convert them to real-space, as detailed in an upcoming article.

Finally, we also note that, beyond the results presented here, it should also be possible to achieve further reduction of the stochastic error with an embedded fragment approach, analogous to self-energy embedding approaches, where a deterministic self-energy is calculated for embedded saturated fragments as introduced for stochastic DFT applications.^{19,20}

AUTHOR INFORMATION

Corresponding Authors

*E-mail: dxn@ucla.edu.

*E-mail: roi.baer@huji.ac.il.

*E-mail: dominika.zgid@gmail.com.

ORCID

Daniel Neuhauser: 0000-0003-3160-386X

Roi Baer: 0000-0001-8432-1925

Notes

The authors declare no competing financial interest.

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ADDITIONAL NOTE

^aA bias arises whenever we plug a random variable x , having an expected value μ and variance σ^2 , into a nonlinear function $f(x)$. One cannot hope that $f(x)$ will have the expected value of $f(\mu)$ unless f is a linear function. A simple example is $f(x) = x^2$, where from the definition of variance $\langle f(x) \rangle = f(\mu) + \sigma^2$. Using the Taylor expansion of f around μ , it is straightforward to show that $f(\bar{x})$, where $\bar{x} = \frac{1}{I} \sum_{i=1}^I x_i$ is an average over I samples and

when I is sufficiently large, $\langle f(\bar{x}) \rangle \approx f(\mu) + \frac{f''(\bar{x})\sigma^2}{2I}$ and so the bias is proportional to the variance of x , the curvature of f at μ and inversely proportional to the number of iterations I .

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